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Unpolarized Gluon Distribution Functions in Leading Order and Next-to-Leading Order

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The unpolarized gluon distribution functions have been obtained by solving Dokshitzer-Gribove-Lipatov-Alterelli-Parisi (DGLAP) evolution equations in LO and NLO at the small-x limit. Here we have used a Taylor series expansion and then the method of characteristics to solve the evolution equations. We have also calculated t and x-evolutions of gluon distribution functions and the results are compared with GRV1998 [1] and MRST2004 [2] gluon parameterizations.

I. THEORY

The DGLAP evolution equations in standard forms for unpolarized gluon distribution functions in LO and NLO [3-10] are

$$\frac{\partial G(x,t)}{\partial t} - \frac{\alpha_{\rm S}}{2\pi} \frac{2}{3} \left\{ \left(\frac{11}{12} - \frac{N_{\rm f}}{18} + \ln(1-x) \right) G(x,t) + I_{\rm g}^{\rm I} \right\} = 0, \tag{1}$$

$$\frac{\partial G(x,t)}{\partial t} - \frac{\alpha_{\rm S}}{2\pi} \frac{2}{3} \left\{ \left(\frac{11}{12} - \frac{N_{\rm f}}{18} + \ln(1-x) \right) G(x,t) + I_{\rm g}^{1} \right\} - \left(\frac{\alpha_{\rm S}}{2\pi} \right)^{2} I_{\rm g}^{2} = 0, \tag{2}$$

where functions $I_g^1(x,t)$ and $I_g^2(x,t)$ are defined in Appendix F.

Let us introduce the variable $u = 1 \square \omega$ and using Taylor's expansion series we can rewrite

$$G\left(\frac{x}{\omega},t\right) = G(x,t) + \frac{xu}{1-u} \frac{\partial G(x,t)}{\partial x} + \frac{1}{2} \left(\frac{xu}{1-u}\right)^2 \frac{\partial^2 G(x,t)}{\partial x^2} + \cdots$$

$$\approx G(x,t) + \frac{xu}{1-u} \frac{\partial G(x,t)}{\partial x}, \qquad (3a)$$

$$F_2^{S}\left(\frac{x}{\omega},t\right) = F_2^{S}(x,t) + \frac{xu}{1-u} \frac{\partial F_2^{S}(x,t)}{\partial x}. \qquad (3b)$$

Since x is small in our region of discussion, the terms containing x^2 and higher powers of x are neglected. Using equations (3a) and (3b) and performing u-integrations we get equation (1) of the form

$$\frac{\partial G(x,t)}{\partial t} - \frac{A_{f}}{t} \left[A_{1}^{g}(x)G(x,t) + A_{2}^{g}(x)\frac{\partial G(x,t)}{\partial x} + A_{3}^{g}(x)F_{2}^{S}(x,t) + A_{4}^{g}(x)\frac{\partial F_{2}^{S}(x,t)}{\partial x} \right] = 0,$$
(4)

where $A_f = \frac{\alpha_s(t)}{3\pi}t = \frac{4}{3\beta_0} = \frac{4}{33 - 2N_f}$, as defined in chapter 3, and also

$$A_{1}^{g}(x) = -\frac{11}{6} + 2x - \frac{1}{2}x^{2} + \frac{1}{3}x^{3} - \ln(x), \qquad A_{2}^{g}(x) = 1 + \frac{4}{3}x - 3x^{2} + x^{3} - \frac{1}{4}x^{4} + 2x\ln(x),$$
$$A_{3}^{g}(x) = \frac{2}{9} \left\{ -\frac{3}{2} + 2x - \frac{1}{2}x^{2} - 2\ln(x) \right\}, A_{4}^{g}(x) = \frac{2}{9} \left\{ 2 + \frac{1}{2}x - 3x^{2} + \frac{1}{2}x^{3} + 4x\ln(x) \right\}.$$



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Now let us assume, $F_2^S(x,t) = R(x)G(x,t)$, where R(x) is a suitable function of x or may be a constant. Now equation (8.4) gives

$$-t\frac{\partial G(x,t)}{\partial t} + A_{f}L_{1}^{g}(x)\frac{\partial G(x,t)}{\partial x} + A_{f}M_{1}^{g}(x)G(x,t) = 0, \qquad (5)$$

where

$$L_{1}^{g}(x) = A_{2}^{g}(x) + R(x)A_{4}^{g}(x) \text{ and } M_{1}^{g}(x) = A_{1}^{g}(x) + R(x)A_{3}^{g}(x) + A_{4}^{g}(x)\frac{\partial R(x)}{\partial x}$$

Now let us consider two new variables S and τ instead of x and t, such that

$$\frac{dt}{dS} = -t , \qquad (6a)$$

$$\frac{\mathrm{d}x}{\mathrm{d}S} = \mathbf{A}_{\mathrm{f}} \mathbf{L}_{1}^{\mathrm{g}}(\mathbf{x}). \tag{6b}$$

Putting these in equation (5), we get

$$\frac{\mathrm{d}\mathbf{G}(\mathbf{S},\tau)}{\mathrm{d}\mathbf{S}} + \mathbf{M}_{1}^{\mathrm{g}}(\mathbf{S},\tau)\mathbf{G}(\mathbf{S},\tau) = 0, \qquad (7)$$

where $M_1^g(S, \tau) = A_f M_1^g(x)$. Equations (7) can be solved as,

$$G(S, \tau) = G(0, \tau) \exp\left[-\int_{0}^{S} M_{1}^{g}(S, \tau) dS\right]$$
(8)

For t-evolution, gluon distribution function varies with t remaining x constant. Hence equation (6a) can be used to solve the equation (7). Now we have to replace the co-ordinate system (S, τ) to (x, t), considering when S = 0, t = t₀ and the input function as $G(\tau) = G(x, t_0)$. So the t-evolution of gluon distribution function in LO is given by

$$\mathbf{G}(\mathbf{x},\mathbf{t}) = \mathbf{G}(\mathbf{x},\mathbf{t}_0) \left(\frac{\mathbf{t}}{\mathbf{t}_0}\right)^{\mathbf{A}_{\mathrm{f}}\mathbf{M}_{\mathrm{I}}^{\mathrm{g}}(\mathbf{x})},\tag{9}$$

Using equation (6b) and replacing the co-ordinate system (S, τ) to (x, t), with consideration when $\tau = 0$, $x = x_0$ and the input function as $G(S) = G(x_0, t)$, the x-evolution of gluon distribution function in LO is given by

$$G(\mathbf{x}, \mathbf{t}) = G(\mathbf{x}_0, \mathbf{t}) \exp\left(-\int_{\mathbf{x}_0}^{\mathbf{x}} \frac{\mathbf{M}_1^g(\mathbf{x})}{\mathbf{L}_1^g(\mathbf{x})} d\mathbf{x}\right)$$
(10)

Similarly the t and x-evolution of gluon distribution functions in NLO are given by

$$G(x, t) = G(x_0, t) \left(\frac{t}{t_0}\right)^{A_f \{M_1^g(x) + T_0 M_2^g(x)\}},$$
(11)

$$G(x, t) = G(x_{0}, t) \exp\left(-\int_{x_{0}}^{x} \frac{M_{1}^{g}(x) + T_{0}M_{2}^{g}(x)}{L_{1}^{g}(x) + T_{0}L_{2}^{g}(x)}dx\right),$$
(12)

where

$$L_{2}^{g}(x) = B_{2}^{g}(x) + R(x)B_{4}^{g}(x) \text{ and } M_{2}^{g}(x) = B_{1}^{g}(x) + R(x)B_{3}^{g}(x) + B_{4}^{g}(x)\frac{\partial R(x)}{\partial x}$$

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We also consider $\left(\frac{\alpha_s(t)}{2\pi}\right)^2 = T_0\left(\frac{\alpha_s(t)}{2\pi}\right)$, where T_0 is a numerical parameter, not arbitrary choose but determined by

phenomenological analysis [79, 174]. The other functions are
$$B_1^g(x) = -\frac{52}{3} \ln x$$
, $B_2^g(x) = -\frac{52}{3} (1 - x + x \ln x)$,

$$B_3^g(x) = \int_x^1 A(\omega) d\omega \text{ and } B_4^g(x) = x \int_x^1 \frac{1-\omega}{\omega} A(\omega) d\omega.$$

II. RESULT AND DISCUSSIONS

We have taken the MRST2004 fit to the H1 [53] and ZEUS [54] data with x < 0.01 and $2 \le Q^2 \le 500 \text{ GeV}^2$ for $Q^2 = 100 \text{ GeV}^2$, in which they have taken parametric form for the starting distribution at $Q_0^2 = 1 \text{GeV}^2$ given by

$$xg = A_{g}x^{-\lambda_{g}}(1-x)^{3.7}(1+\varepsilon_{g}\sqrt{x}+\gamma_{g}x) - Ax^{-\delta}(1-x)^{10}$$

where power of the $(1\Box x)$ factors are taken from MRST2001 fit [56]. Here $A_g, \lambda_g, \gamma_g, \varepsilon_g$, A and δ are taken as free parameters. The optimum fit corresponds to $\alpha_s (M_z^2) = 0.119$ and $\Lambda_{\overline{MS}} = 323$ MeV with $N_f = 4$.

Our results represent the best fit graphs of our work with different parameterization curves. Results of parameterization at lowest- Q^2 values are taken as input to test the t-evolution equations and those at highest-x is taken as input to test the x-evolution equations. We have compared our results for R(x) as a constant R, a power function ax^b and an exponential function $ce^{\Box dx}$. In our work for gluon distribution function, we have found the values of the gluon distribution function remains almost same for b<0.0001 and for d<0.001. So, we have chosen b = 0.0001 and d = 0.001 and the best fit graphs are observed by changing the values of R, a and c. If we plot T²(t) and T₀T(t) against Q², then we can see that for T₀ = 0.048, the values of T² and T.T₀ are nearly same in our region of discussion, as we have seen in figure 4.1 of chapter 4. Thus we consider T₀ = 0.048 in calculation of G(x, t) at NLO and the consideration of parameter T₀ does not give any abrupt change in our results.

In figures 8.1, we have plotted our results of t-evolution of gluon distribution function in LO from equation (9) and compared with GRV1998 parameterization for R(x) = R, a constant. Here we have plotted our results of gluon distribution function against Q^2 for $x = 10^{-5}$ and $x = 10^{-4}$ and we get the best fit with R = 0.5 and R = 0.8 respectively.

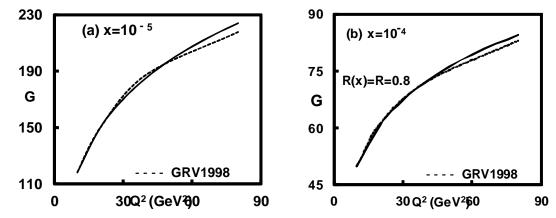


Figure 1: t-evolution of gluon distribution functions in LO for R(x) = R, a constant, compared with GRV1998 parameterization graphs



In figures 2, we have plotted our results of t-evolution of gluon distribution function in LO from equation (9) and compared with GRV1998 parameterization at x=10 ^{\Box 05} for R(x) = ax^b and ce^{\Box dx} respectively. Here we have plotted our results against Q² for R(x) = ax^b as well as R(x) = ce^{\Box dx} and we get the best fit with a = 0.9, b = 0.0001 and c = 0.8, d = 0.001.

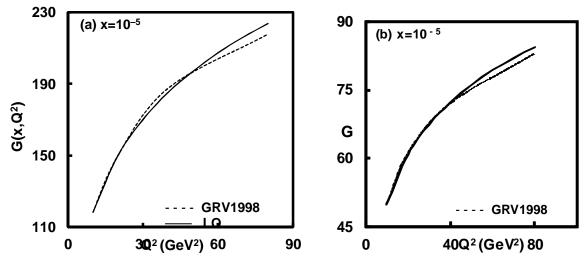


Figure 2: t-evolution of gluon distribution functions in LO for $R(x) = ax^b$ and $ce^{\Box dx}$ compared with GRV1998 parameterization graphs

In figure 3 we have plotted our results of gluon distribution function against x for $Q^2 = 100 \text{ GeV}^2$ with R(x) = R and compared with MRST2004 and best fit has been found for R = 0.4.

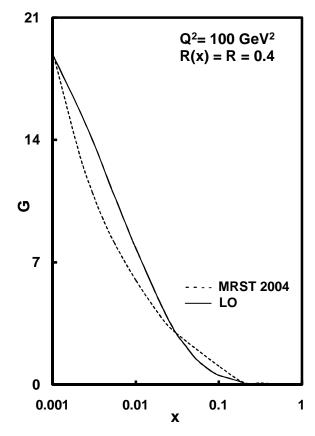


Figure 3: x-evolution of gluon distribution functions in LO for R(x) = R, compared with MRST2004 parameterization



In figures 4 and 5, we have plotted our results of t-evolutions of gluon distribution function in NLO from equation (11) and compared with GRV1998 gluon parameterization for R(x) = R and ax^b . These results are also compared with our LO results obtained from equation (9). In figure 4, we have plotted our results of gluon distribution function against Q² for $x = 10^{-5}$ and we get the best fit with R = 0.8 in NLO and R = 0.9 in LO.

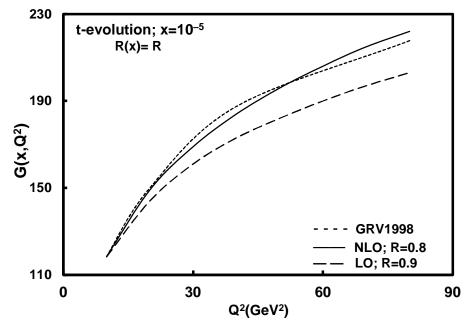


Figure 4: t-evolution of gluon distribution functions in NLO for R(x) = R, compared with GRV1998 parameterization

In figures 5 our results of gluon distribution function in NLO have been plotted against Q^2 for $x = 10^{-4}$ with $R(x) = ax^b$ and we get the best fit with a = 0.8 and b = 0.001 in NLO and a = 0.9 and b = 0.0001 in LO.

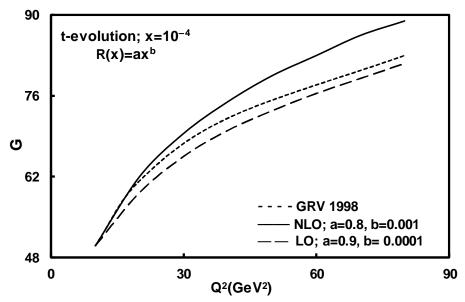


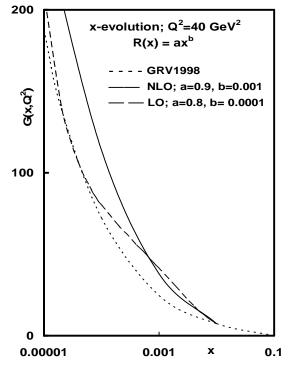
Figure 5: t-evolution of gluon distribution functions in NLO for $R(x) = ax^b$, compared with GRV1998 parameterization

In figure 6, we have plotted our results of x-evolution of gluon distribution function in NLO from equation (12) and compared with GRV1998 gluon parameterization for $R(x) = ax^b$. These results are also compared with our LO results. We have plotted our results for $Q^2 = 40 \text{ GeV}^2$ and the best fit has found for a = 0.9 and b = 0.001 in NLO and for a = 0.8 and b = 0.0001 in LO.



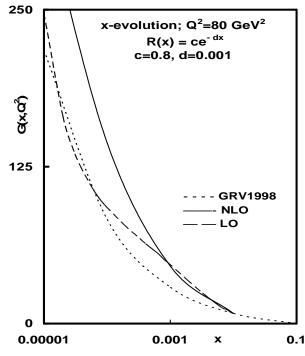
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Figures 6: x-evolution of gluon distribution functions in NLO for $R(x) = ax^b$, compared with GRV1998 parameterization

In figure 7, we have plotted our results of x-evolution of gluon distribution function in NLO from equation (12) and compared with GRV1998 gluon parameterization for $R(x) = ce^{\Box dx}$. These results are also compared with our LO results. We have plotted our results for $Q^2 = 80 \text{ GeV}^2$ and the best fit has found with c = 0.8 and d = 0.001 in both NLO and LO.



Figures 7: x-evolution of gluon distribution functions in NLO for $R(x) = ce^{-dx}$, compared with GRV1998 parameterization



In figure 8 we have plotted our results of gluon distribution function against x for $Q^2=100 \text{ GeV}^2$ with R(x) = R, a constant and $R(x) = ax^b$, a power function of x. Our results are compared with MRST2004 as well as our LO results. The best fit has been found with R= 0.4 in both NLO and LO and also with a = 0.5, b = 0.0001 in NLO and a = 0.5, b = 0.001 in LO respectively.

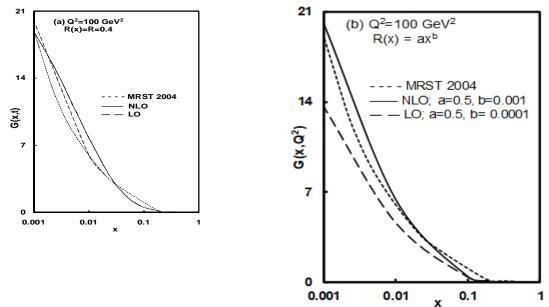


Figure 8: x-evolution of gluon distribution functions in NLO for R(x) = R, a constant and ax^{b} , compared with MRST2004 parameterization

III. CONCLUSION

In this chapter, we have solved the DGLAP evolution equation by using method of characteristics and get gluon distribution function in LO and NLO. We have calculated here the t and x-evolutions of gluon distribution functions. It is shown that our results are in good agreement with GRV1998 and MRST2004 global parameterizations especially at small-x and high- Q^2 region. Here from global parameterizations and our results we have seen that the gluon distribution functions increase when x decreases and Q^2 increases for fixed values of Q^2 and x respectively. On an average, the mean percentage errors of our LO and NLO results are 7.73% and 1.86% with GRV1998 and also 11.45% and 3.63% with MRST2004 global parameterizations respectively. These errors of our results are very less as compared to systematic and statistical uncertainties in the experimental data. Thus, there is significant contribution of Next–to–Leading order over the Leading order in unpolarized gluon distribution functions.

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