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Review on Automatic Antenna Tracking System For LEO Satellites

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Abstract: This paper is about comparing the Yagi- Uda Antenna, Turnstile Antenna, Parabolic Antenna Phased Array Antenna with a different parameters required to track LEO Satellites and to receive the Telemetry Information from them. LEO Satellites is Low Earth Orbit typically organized as a satellite constellation. The number of Satellite LEO would be ten to even thousand to fully cover the globe. As LEO satellites move very quickly and are most visible for 20 to 30 min during each pass, it requires an antenna that can track signals, and satellite paths, and upload and download as much data as possible in a short amount of time. The continual motion of tracing one LEO satellite after other equates to significant mechanical performance. So in this paper, we carry out a review of which antenna suites are best for tracking the LEO, getting data from them, and also the mechanical parameter of Antennas.

Keywords: LEO satellite, Constellation.

I. INTRODUCTION

THE BEGINNING OF THE NEW SATELLITE ERA

Since the beginning of new the millennium, the whole aerospace community has evolved rapidly into a new era by revolutionizing satellite technologies [1]-[3]. Traditional satellites were bulkier and more expensive but due to the advancement in technologies, the new satellites are of advanced technology less expensive and a few kilograms, and also smaller in size. These small satellites are often classified according to their weights i.e mini (100- 500kg), Micro (10-100kg), Nano (1-10kg), Pico (0.1-1kg), Femto.1kg) [3]. A well-known such small satellite is called CubeSat, which is made of several units with a fundamental size of 10 x10 x 10 cm³ cube (1U). For their low size and weight, small satellites have attracted not only traditional institutional players such as government, universities, and space agencies but also intensive investment from private companies.

Unlike the classical single big satellite in geostationary orbit (GEO) and the medium earth orbit (MEO), small satellites in low earth orbit (LEO) are typically organized as a satellite constellation. The (LEO) satellite constellation would be several hundred or even ten thousand to fully global coverage. A Group of satellites is deployed at the same altitude and inclination in several orbital planes. There are lots of advantages to LEO satellite constellation, such as small propagation loss, high fault tolerance as well as network robustness. With the low orbit of LEO satellites, the one-way propagation delay is only up to 15ms, which is commensurate with that of terrestrial links. Meanwhile, much less power is required to overcome propagation loss and other atmospheric attenuation and the network structure of the LEO satellite constellation also provides resilience as a space network.

The concept of the LEO satellite constellation has the potential to open exciting innovation opportunities for space exploration and lead to exciting scientific breakthroughs. Table 1 list some of the recent LEO satellite constellation.

Moreover, there is a trend of new LEO satellite constellation projects for global Internet access. Interestingly, most of them are invested by giant Internet companies [16]. For example, SpaceX conducts the manufacturing of the satellites and launches them using their own Falcon vehicles. As of July 2021, 1738 satellites of the Starlink constellation have been deployed. and command (TT&C) service. It should drive the antenna beam to follow the satellite during the interval from its rise from the horizon until it sets below the horizon.

The period of a pass is relatively short for an LEO satellite, with a typical duration of a few minutes, depending on the relative position between the satellite and the ground station. Traditionally, rector antennas are used as satellite ground stations with high performance and medium cost [17], [18]. However, it can only support a limited number of satellites in its very narrow beam. This problem can be alleviated by increasing the number of rector antennas at the mission control site. However, the total cost of land, infrastructure, and antennas forbids such a mode of antenna farms.

Hence, rector antennas cannot meet the requirements of ordinary management and control operation of the future ultra-dense LEO satellite constellation. To solve the aforementioned problem, phased array antennas are considered attractive candidates for the future ground station of the LEO satellite constellation.

The primary reason is that phased array antennas can produce a number of beams by the beamforming network thus supporting multiple satellites simultaneously [19]-[21]. There are also other advantages including faster beam steering, better interference mitigation, higher reliability, and lower cost (see Section 2 for detailed discussion). Previously the cost of phased array antennas are very high, thus early applications were coned to military and defense areas.

II. CHARACTERISTICS OF ANTENNAS

The performance of the antenna, which includes the description of how the radiating waves behave in space and in relation to the antenna, is described by several parameters.

A. Directivity

Yagi-UDA antenna has high directivity by applying current phase shifts between the elements due to resonance phenomena. There are directors and reflectors which are nonelectric elements. The director helps to improve the directivity of the yagi-UDA antenna.

Turnstile antenna provides high directivity, several turnstiles may be piled up along a vertical axis.

The main advantage of a parabolic antenna, it has high directivity. It functions similarly to a flashlight reflector to direct radio waves in thin beam radio waves from one specific direction only.

Phased array antenna Directivity is the ratio of radiation intensity in a given direction from an antenna to the radiation intensity averaged over all directions. If that specific direction is not underlined, then the direction in which max intensity is observed can be taken as the directivity.

B. Gain

Yagi UDA antenna The structure of this directional antenna integrates three main parts which are directors, reflectors, and a feeder point of 50Ω . The performance of the proposed antenna has a tolerable reflection coefficient below -10 dB at 144.46 MHz. A maximum gain of around 14.03 dB with a good efficiency of 75.6% is achieved.

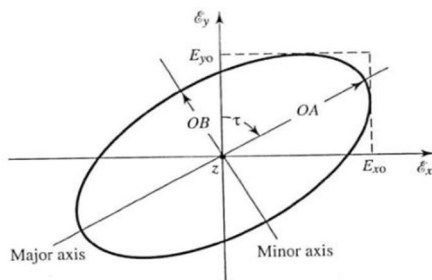
Turnstile antenna the gain of a turnstile antenna is actually about 3 dB less than that of a single dipole in its direction of maximum radiation because each of the elements of the turnstile antenna receives only one-half the tx power.

Parabolic antenna gain is calculated as the gain over an isotropic source, i.e. radiates equally in all directions.

Phased antenna with a very large number of regularly- spaced radiating elements, the gain realized at the beam peak is equal to the number of elements times the gain realized in the same direction when only one element is excited. The ideal radiation pattern of one such element in a large planar array has a cosine variation of gain with an angle when the elements are closely spaced, and has a peak value of gain equal to $4\pi A/\lambda^2$ where A is the area allotted to each element. The active impedance of each element in a practical phased array varies with the scan angle, because of the mutual coupling between the elements. The associated mismatch causes power to be returned to the generators, thereby reducing the gain realized by the array and by the element.

C. Polarization

Since these three interpretations are related, the interpretation of polarization for a vector at some point in space and for a plane wave is part of the explanation of polarization of an antenna. The polarization to a vector is the shape, orientation, and sense of the ellipse that the extremity of the vector describes as a function of time. For polarization of a plane wave, the following applies, In a single-frequency plane wave a vector has the same polarization at every point in space. Lastly, for an antenna, the term polarization is used as follows. The polarization of an antenna in a given direction of the plane wave it radiates at huge distances in that direction. The polarization can be described as linear, circular, or elliptical. Linear and circular polarization can be seen as of elliptical polarization. The linear case is when the electrical vector at a point in space is a function of time points along a line.



The Yagi Uda antenna can radiate waves as both horizontal and vertical signals.

Turnstile

Two new turnstile antennas are described: 1) a horizontally polarized turnstile antenna and 2) a circularly polarized turnstile antenna. These antennas feature a rotating phase feed based on the original RCA Super turnstile design which, because of its simplicity, electrical characteristics, and versatility, is today the most popular TV broadcast antenna.

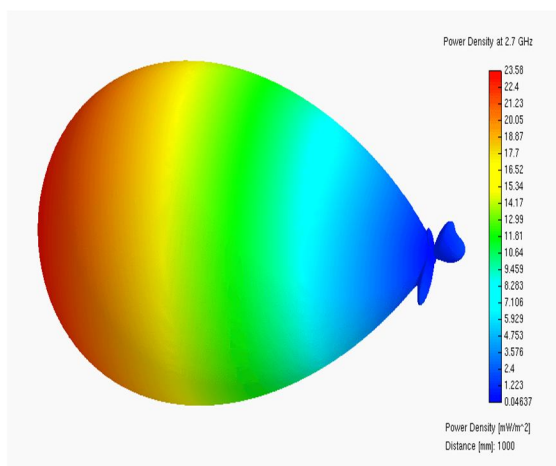
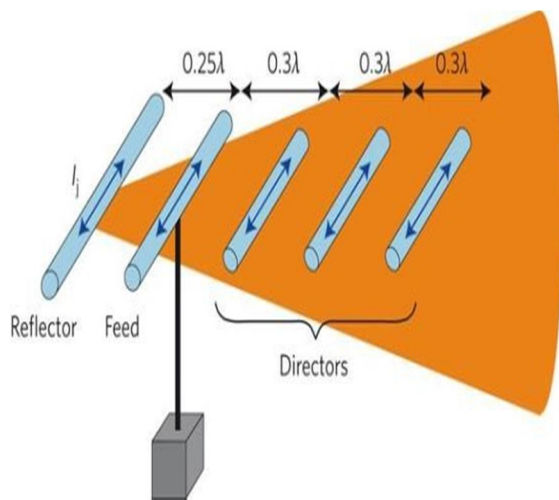
Parabolic antenna polarization. The pattern of electric and magnetic fields at the front of a parabolic antenna is a scaled-up image of the radiation by the feed antenna, so the polarization is determined by the feed antenna.

Phased array Dual Polarized (DP) Phased Array Radar (PAR) performance and cost are greatly influenced by the radiating elements selected and the architecture of the module. The performance can be dynamically optimized for a particular scan angle using prior knowledge of the characteristics of the radiating elements. The effectiveness of this strategy and the degree of difficulty in implementing it depends on major system decisions regarding scan volume, target objective (minimum detectable signal), and cost.

D. Radiation Pattern

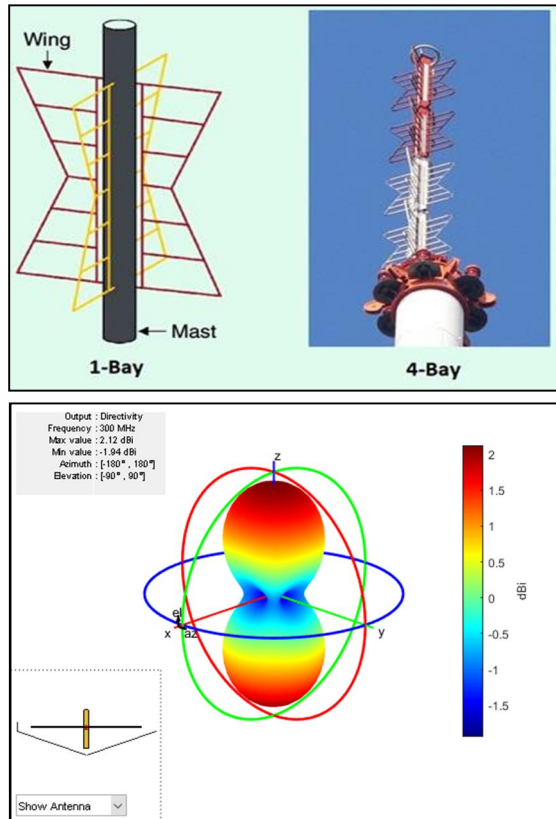
When it comes to satellite communication, the ground station and the satellite are definitely in each other's far field. However, when antenna measurements are done at ground level if, for instance, a helical antenna were to be tested, the distance between the antenna and the receiver must be selected in such a manner that the measurements are done in the far field region. The radiation pattern consists of lobes of various sizes, the direction with the maximum radiation intensity is known as the major lobe or the beam of the antenna. The half-power beam width (HPBW) is measured. The HPBW is the angle between the two directions where radiation intensity has decreased to half of the maximum radiation. This angle is inversely proportional to the square of directivity or the gain and is a useful parameter when the gain is difficult to measure.

Yagi Uda



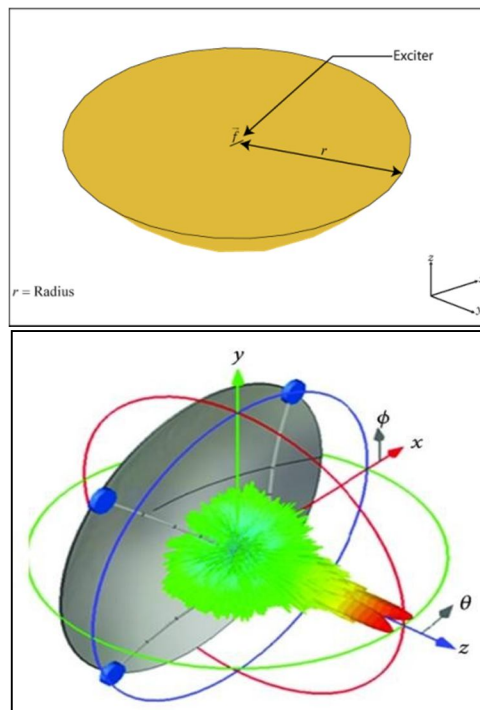
Yagi-Uda antenna have high directivity as shown in the figure given below. The minor lobes are reduced and the directivity of the major lobe is improved by the addition of directors.

Turnstile antenna



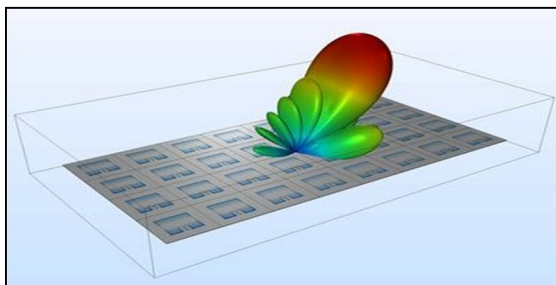
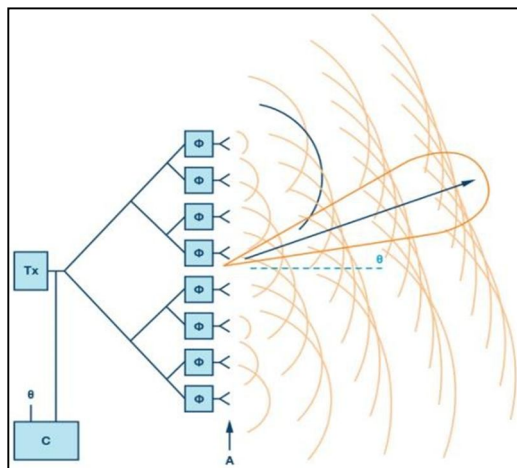
The radiation pattern of a parabolic antenna contains a major lobe, which is directed along the axis of propagation, and several small minor lobes. Very thin beams are developed with this type of reflector.

Phased antenna



The radiation pattern will be similar to the radiation pattern of two superimposed dipoles. Though it is close to an omnidirectional pattern, it leaves a cloverleaf-shaped pattern.

Parabolic antenna



each element produces an equal output voltage of $\sin(\omega t)\sin(\omega t)$. However, there will be an equal phase difference Ψ between successive elements. Mathematically, it can be written as $\Psi = 2\pi d \sin \theta / \lambda$ Equation 1 Where, θ is the angle at which the incoming signal is incident on each radiation element.

E. Mechanical strength.

Yagi UDA antennae are easy to install on the rotor mechanism and due to being lightweight it has low wind resistance and turning also required less power.

The turnstile antenna is easy to install on the rotor but it has a little more wind resistance compared to the yagi UDA antenna.

Parabolic antennae required a beam focused in the direction of reception of satellite signal so it will be difficult to accurately point it towards the satellite as the size increases the rotor cost also increases and it will be difficult to rotate.

Phased antennae are grp of antenna which is organized in a pattern. As they are grp it has high wind resistance and rotation due to this load requiring a high-duty cycle. So it is not convenient to use this antenna.

III. CONCLUSION

To meet the requirements of ordinary management and control operation of future ultra-dense LEO satellite constellations, Yagi uda antennas are considered as attractive candidates thus supporting multiple satellites simultaneously. Many advantages of yagi uda antennas have been recognized compared to traditional reflector antennas in satellite ground stations, including better performance, higher reliability, and lower cost.

IV. FUTURE SCOPE

We are planning to make a yagi uda antenna as per all the comparisons made, yagi UDA is reliable and best to use for automated tracking of Leo satellite, and thus in the future, we will make an AQQ controller which will require telemetric data of satellite and it will track the Leo constellation simultaneously and result would be displayed



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